

MATHEMATICAL TRIPOS Part III

Thursday 9 June, 2005 1.30 to 3.30

PAPER 76

DYNAMO THEORY

*Attempt **TWO** questions.*

*There are **THREE** questions in total.*

The questions carry equal weight.

STATIONERY REQUIREMENTS

*Cover sheet
Treasury Tag
Script paper*

SPECIAL REQUIREMENTS

None

**You may not start to read the questions
printed on the subsequent pages until
instructed to do so by the Invigilator.**

1 (i) A uniform magnetic field \mathbf{B}_0 permeates a fluid of magnetic diffusivity η and is subjected to the effect of the differential rotation associated with the velocity field $\mathbf{u} = \boldsymbol{\omega}(\mathbf{x}) \times \mathbf{x}$, where $\boldsymbol{\omega}(\mathbf{x})$ is everywhere parallel to \mathbf{B}_0 . Using cylindrical polar coordinates (s, ϕ, z) , with Oz parallel to \mathbf{B}_0 , so that $\mathbf{B}_0 = (0, 0, B_0)$ and $\mathbf{u} = (0, s\omega(s, z), 0)$, show that a toroidal magnetic field $(0, B(s, z, t), 0)$ is generated, where B satisfies the equation

$$\frac{\partial B}{\partial t} = s(\mathbf{B}_0 \cdot \nabla)\omega + \eta \left(\nabla^2 - \frac{1}{s^2} \right) B. \quad (*)$$

(ii) Now adopt spherical polar coordinates (r, θ, ϕ) and suppose that $\omega = \omega(r)$ where $\omega(r)$ is a smooth function of r , finite at $r = 0$ and falling to zero more rapidly than r^{-6} as $r \rightarrow \infty$. Verify that equation $(*)$ admits a steady solution of the form

$$B = -\frac{B_0}{\eta} \frac{\sin \theta \cos \theta}{r^3} \int_0^r x^4 \omega(x) dx.$$

(iii) Estimate the time it would take to establish this steady solution in a fluid of very high conductivity, starting from an initial condition $B = 0$ at $t = 0$.

(iv) Explain in physical terms the effect of the same differential rotation if \mathbf{B}_0 is in the plane of the velocity field \mathbf{u} (rather than perpendicular to this plane).

2 (i) Explain the principles of *Mean Field Electrodynamics* leading to the equation

$$\boldsymbol{\mathcal{E}} = \alpha \mathbf{B} - \beta \nabla \times \mathbf{B} + \dots,$$

where $\boldsymbol{\mathcal{E}}$ is the mean electromotive force associated with a homogeneous isotropic field of turbulence, and \mathbf{B} is the mean magnetic field.

(ii) Using the first-order smoothing approximation (which should be justified), obtain a relationship between the parameter α and the mean helicity of the turbulence.

(iii) Obtain a criterion for dynamo action resulting solely from the α -effect.

(iv) Discuss the application of this theory to the problem of the geodynamo.

3 Consider an interface dynamo with a magnetic field

$$\mathbf{B}(x, z, t) = \begin{cases} (-\partial A/\partial z, B, \partial A/\partial x) & z > 0 \\ (-\partial a/\partial z, b, \partial a/\partial x) & z < 0 \end{cases}$$

referred to Cartesian co-ordinates, such that the components of \mathbf{B} satisfy the scalar equation

$$\frac{\partial B}{\partial t} - \eta \nabla^2 B = 0 \quad , \quad \frac{\partial A}{\partial t} - \eta \nabla^2 A = \alpha B$$

for $z > 0$, and the equation

$$\frac{\partial a}{\partial t} - \eta \nabla^2 a = 0 \quad , \quad \frac{\partial b}{\partial t} - \eta \nabla^2 b = V \frac{\partial a}{\partial x}$$

for $z < 0$, where η, α, V are positive constants. Explain the significance of the various terms in these equations and obtain the four conditions relating A, a and B, b at $z = 0$, given that the mean velocity is continuous. [Hint: remember the electric field.]

Now consider surface dynamo waves with

$$\begin{aligned} (A, B) &= (A_0 + Cz, B_0) e^{st - \Lambda z} e^{i(kx - Lz - \omega t)}, \\ (a, b) &= (a_0, b_0 + cz) e^{st + \lambda z} e^{i(kx + lz - \omega t)}, \end{aligned}$$

where Λ, λ are positive constants. Show that the amplitude of these waves grows exponentially with time if $D > 32$, where the dynamo number

$$D = \frac{V\alpha}{\eta^2 k^3}.$$

How can this simple model be related to cyclic activity in stars like the Sun?

END OF PAPER